

# Whole Atmosphere Community Climate Model-eXtended (WACCM-X):

sun

# Development, Validation and Capabilities

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CEDAR Prize Lecture, NSF CEDAR Workshop, June 28, 2018, Santa Fe, NM



earth•connections

# Outline

- WACCM-X Development Background.
- Recent Development and Validation.
- Model Capabilities for Research.
- Ongoing and Future Efforts.

### Genesis—It goes back in TIME (GCM):

Thermosphere-Ionosphere-Mesosphere-Electrodynamics General Circulation Model

- Developed by Ray Roble and colleagues since the 1970s.
- Validated and widely used by the CEDAR community.
- 1) 1989, Arthur Richmond (HAO/NCAR) Assimilative Mapping of Ionospheric Electrodynamics
- 2) 1990, Michael Mendillo (Boston U) The Discovery of a Sodium Magneto-Nebula Around Jupiter
- 3) 1991, Craig Heinselmann (SRI International) Sondrestrom MUSCOX
- 4) 1992, Colin Hines (Arecibo Obs) The Doppler Spreading Theory of Gravity Wave Spectra
- 5) 1993, John Cho (Arecibo Obs), Radar Scattering from the Coldest Place in our Atmosphere: Polar Mesosphere Summer Echoes
- 6) 1994, Raymond Roble (HAO/NCAR), Modelling the Circulation, Temperature and Compositional Structure of the Upper Atmosphere (30-500km)
- 7)1995, David Fritts (U of Colorado) Modeling of Gravity Wave and Instability Processes in the Middle Atmosphere
- 8) 1996, Chester Gardner (U of Illinois) The ALOHA/ANLC-93 Campaigns
- 9) 1997, Bela Fejer (Utah State U) Multi-Instrument Studies of Ionospheric Electrodynamics

### From Ray's Prize Lecture in 1994

#### LONG RANGE TGCM MODEL DEVELOPMENT

#### MODEL INPUTS PROGRAMS 1979 Ionosphere Atmosphere Explorer Dynamo TGCM Dynamics Explorer MSIS (95 - 500 km) Radar and Airglow $\rightarrow$ CEDAR Aurora Solar Tides 1989 Dynamo Atmosphere Explorer TIGCM Aurora Dynamics Explorer $\rightarrow$ Solar (95 - 500 km) Radar and Airglow Tides CEDAR Air Force 1991 Aurora (AMIE) Atmosphere Explorer TIE-GCM Solar Dynamics Explorer $\rightarrow$ Tides (95 - 800 km) CEDAR GEM MTIE-GCM ISTP/GGS 1992 Aurora (AMIE) UARS Solar (UARS) TIME-GCM Solar Mesosphere Explorer $\rightarrow$ Tides and Waves (GSWM) (30 - 500 km) CEDAR GEM Air Force TIMED





# Further Development to Better Represent the Lower Atmosphere

- From climatology to climate/weather variability:
  - Atmospheric tidal climatology, specified by the Global Scale Wave Model (GSWM).
  - Gravity wave drag parameterization.
  - Lower atmosphere by NCEP or ECMWF data.
- Self-consistent representation of the whole atmosphere system
  - Coupled TIME-GCM/CCM3.
  - WACCM/WACCM-X.

Scientific Objectives of Whole Atmosphere Community Climate Model and its extended version (WACCM/WACCM-X)

- Solar impacts on the Earth System.
- Understand and quantify couplings between atmospheric layers through chemical, physical and dynamical processes.
- Implications of the couplings to climate (downward coupling) and to space environment (upward coupling).

### NCAR Community Earth System Model (CESM)





### Major CESM WACCM/WACCM-X Components

Model Framework	Chemistry	Physics	Physics	Resolution
Atmosphere component of NCAR Community Earth System Model (CESM) Extension of the NCAR Community Atmosphere Model (CAM) Finite Volume Dynamical Core (modified to consider species dependent Cp, R, m) Spectral Element Dynamical Core	MOZART+ lon Chemistry (~60+ species) Fully-interactive with dynamics.	Long wave/short wave/EUV RRTMG IR cooling (LTE/non- LTE) Modal Aerosal CARMA Convection, precip., and cloud param. Parameterized GW Major/minor species diffusion (+UBC) Molecular viscosity and thermal conductivity (+UBC) Species dependent Cp, R, m.	Parameterized electric field at high, mid, low latitudes. IGRF geomagnetic field. Auroral processes, ion drag and Joule heating Ion/electron energy equations Ambipolar diffusion Ion/electron transport Ionospheric dynamo Coupling with plasmasphere/ magnetosphere	Horizontal: 1.9° x 2.5° (lat x lon configurable as needed) Vertical: 66 levels (0-140km) 81/126 levels 0-~600km Mesoscale- resolving version: 0.25 deg/0.1 scale height.

### **Important Physical Properties of Whole Atmosphere**

- Deep: 10% of Earth radius, ~29 scale-heights, 10<sup>13</sup> change in density from Earth surface to exobase.
- Diffusive separation above the homopause.
- Ion-neutral coupling
  - Different transport processes of neutral species and ionospheric plasma (oriented along magnetic field lines).
- Coupling between dynamics and photochemistry.
- Short temporal and spatial scales in the upper atmosphere
  - Increasing significance of gravity waves and tides.
  - Large wind (~300m/s) and acoustic speed (~800m/s).
  - Geomagnetic storms and fine ionospheric structures.
  - Ionospheric irregularities.
  - Large molecular viscosity and diffusion (both vertically and horizontally).

Implications for Mathematical and Numerical Formulation (1)

- Diffusive separation above the homopause:
  - Specific heats and mean molecular weight (thus gas "constant" of dry air) are dependent on major species (O, O<sub>2</sub>, N<sub>2</sub>, He, H), thus vary spatially and temporally.
  - Potential temperature becomes an ill-posed quantity: the mixing ratios of the major species are different from those at reference levels.
  - Variable gravity affects the scale height (thus the vertical distribution) of individual species.



Horizontal winds and divergence are solved incorrectly (and often become too strong) with the standard formulation. Causes excessive upwelling in the summer and downwelling in the winter.







With k advection

### Implications for Mathematical and Numerical Formulation (2)

- Coupling between dynamics and photochemistry.
  - Conservative and efficient computation of advective transport of large number of chemical species.
- Ion-neutral coupling:
  - Frequent mapping between dycore grid and geomagnetic grid.
  - Transport routines that can handle different advective velocities (neutral winds and ion velocities).

Implications for Mathematical and Numerical Formulation (3)

- Short temporal/spatial scales of physical processes:
  - Parameterization schemes that can accommodate short time steps (5 minutes or less).
  - Code design that is capable of subcycling and supercycling.
  - Mesh refinement capability.
  - Non-hydrostatic dynamics.
  - Horizontal diffusion should be included with increasing spatial resolution.
  - Efficient scaling for high-resolution simulations.

### **Ionospheric Electric Dynamo**

Ionospheric electrostatic potential is solved by using Ohm's Law and current continuity condition (Richmond, 1983)

 $\nabla \bullet (\sigma : \nabla \Phi) = \nabla \bullet (\sigma : (\overrightarrow{V} \times \overrightarrow{B})) + \text{Highlatitude electric potential}$ 



Fig. 6. Block diagram connecting the physical attributes at work in the E- and F-region dynamos.

Heelis, 2004 (CEDAR Tutorial)

# F-region O<sup>+</sup> Transport and Electron/Ion Temperatures

- O+ transport determined by field aligned ambipolar diffusion and ExB drifts.
- Ambipolar diffusion depends on electron and ion temperatures.
- Te tendency considered: vertical component of electron heat conduction along field-line and heating/cooling.
- Heating of neutrals by thermal electrons and ions are now included in the model.

### **Key WACCM-X Capabilities**

- Physics-based whole atmosphere general circulation model (0-700km)
- Solves dynamics, radiative transfer, photolysis and energetics
- Fully interactive chemistry, including ion chemistry.
- Ionospheric electrodynamics using fully interactive dynamo.
- Ion transport in the *F*-region.
- Magnetospheric inputs using empirical or specifications, including AMIE.
- Coupling with a plasmasphere model (partnership with NRL).
- Meteorology can be constrained by reanalysis data (MERRA).
- Whole atmosphere data assimilation for specification and forecast.
  - WACCM-X Tutorial during 2017 CEDAR Workshop
  - WACCM-X has been released as part of CESM2.



### Mass and Electron Density at 400km









# J. Liu et al. 2018





J. Liu et al. 2018

### **NmF2 From WACCM-X**



### 2003 Halloween Storm Simulated by WACCM-X



TEC, SD-WACCM-X, Hourly Time: 2003-10-19 02:00

Courtesy of Chuck Bardeen

### WACCM-X has also been used for:

#### AGU100 ADVANCING EARTH AND SPACE SCIENCE



#### **Geophysical Research Letters**

#### **RESEARCH LETTER** 10.1029/2018GL077723

solar eclipse effects on geo-

space: The 21 August 2017

New understanding of the

Special Section:

Solar Eclipse

Simulation of the 21 August 2017 Solar Eclipse Using the Whole Atmosphere Community **Climate Model-eXtended** 

Joseph M. McInerney<sup>1</sup> (0), Daniel R. Marsh<sup>1,2</sup> (0), Han-Li Liu<sup>1</sup> (0), Stanley C. Solomon<sup>1</sup> (0), Andrew J. Conley<sup>2</sup> (D), and Douglas P. Drob<sup>3</sup> (D)





#### **Geophysical Research Letters**

**RESEARCH LETTER** 

10.1029/2018GL077867

Key Points:

**Kev Points:** 

The Influence of Internal Atmospheric Variability on the lonosphere Response to a Geomagnetic Storm

 WACCM-X simulations demonstrate the impact of lower atmosphere variability during a geomagnetic storm

N. M. Pedatella<sup>1,2</sup> and H.-L. Liu<sup>1</sup>

#### **AGU** PUBLICATIONS



#### Journal of Geophysical Research: Space Physics

**RESEARCH ARTICLE** 10.1002/2017JA024998

**Temporal Variability of Atomic Hydrogen From** the Mesopause to the Upper Thermosphere

**Key Points:** 

· In the mesopause, hydrogen density is higher in summer than in winter, and higher at solar minimum than at

Liying Qian<sup>1</sup> (0), Alan G. Burns<sup>1</sup> (0), Stan S. Solomon<sup>1</sup> (0), Anne K. Smith<sup>2</sup> (0), Joseph M. McInerney<sup>1</sup> (0), Linda A. Hunt<sup>3</sup> [0, Daniel R. Marsh<sup>1,2</sup> [0, Hanli Liu<sup>1</sup> [0, Martin G. Mlynczak<sup>3</sup> [0, and Francis M. Vitt<sup>1,2</sup> [0]

#### **@AGU** PUBLICATIONS



#### **Geophysical Research Letters**

**RESEARCH LETTER** 10.1002/2017GL076950

comprehensive whole-atmosphere

· We have performed the first

Whole Atmosphere Simulation of Anthropogenic **Climate Change** 

Stanley C. Solomon<sup>1</sup>, Han-Li Liu<sup>1</sup>, Daniel R. Marsh<sup>1</sup>, Joseph M. McInerney<sup>1</sup>, Liying Qian<sup>1</sup> (), and Francis M. Vitt<sup>1</sup>

climate change simulations, including

## **Equatorial Ionosphere Weather**

### **ExB Drifts: WACCM-X vs Climatology**



Liu et al., 2018

Dotted line: JRO climatology (Fejer et al., 1991)

### Monthly Mean PRE Peak



### Monthly vs Daily Variability





Line contour: Cor(PRE,U\*Ped(19LT)); color: StdDev(U\*Ped(19LT))

- Summer side E-region neutral wind variability at sunset time strongly affects PRE variability.
- E-region is strongly affected by lower atmospheric waves.



### **Occurrence Frequency of Equatorial Plasma Bubbles**

100

80

60

40

20





- The agreement between deduced EPB rates and the observed rates suggest
  - Large-scale dynamics and electrodynamics play a key role in preconditioning EPB
  - Feasibility for probabilistic forecast of EPB—an outlook/warning (analogous to tornado forecast).
- Resolving EPB requires high-resolution capability.



### WACCM-X Data Assimilation: WACCM-X+DART

lonosphere results based on assimilating data only below 100km (meteorological observations, Aura/MLS, and TIMED/SABER): Large-scale ionospheric features may be forecast 10-20 days in advance.



Pedatella et al., 2018



### Model Biases and Uncertainties: Mixing by Subgrid Scale Waves



## Model Biases and Uncertainties: Forcing by Subgrid Scale Waves



Pedatella et al. (2014)

# Model Biases and Uncertainties

- A crucial missing physics is the gravity waves
  - Gravity waves are the major driver of upper atmosphere dynamics.
  - Not well represented due to insufficient spatial resolution.
  - Usually parameterized: a major source model bias and leads to predictive errors.
- Need to:
  - resolve them, and/or
  - Better parameterization and constraints, and/or
  - perform whole atmosphere data assimilation.

# Summary

- Key WACCM-X capabilities have been developed, and validated against thermospheric and ionospheric observations for climatology, variability during geomagnetic quiet and disturbed conditions, and longterm space climate change.
- Simulated PRE, an important quantity for the formation of EPB, shows longitudinal and seasonal variation similar to observations.
  - Simulated PRE varies significantly from day-to-day. Deduced EPB rate is similar to observations.
- Model biases in mesosphere, thermosphere and ionosphere can be caused by issues with gravity wave parameterization (both drag and mixing) (déjà vu...)

# CESM2/WACCM-X v2: A Community Model

CESM2/WACCM-X is a community model—play with it!

• WACCM-X Link:

https://www2.hao.ucar.edu/modeling/waccm-x

• CESM Link:

http://www.cesm.ucar.edu/

### Ongoing/Future WACCM-X Developments

- Perform WACCM-X simulations in support of upcoming satellite missions (ICON, GOLD, and COSMIC-2) as well as ground-based observations, including their use in data assimilation.
  - Can be a powerful tool integrating and interpreting observations.
- Develop generic 3D mapping capability between new WACCM-X dynamical core grids and geomagnetic grid.
- Develop WACCM-X with mesoscale-resolving capability.
- Assimilation of ionospheric observations.
- Including Helium as a major species.
- Ionospheric E-region transport and metal ion chemistry.
- D-region chemistry.
- Whole geospace modeling by coupling WACCM-X to magnetosphere and plasmasphere models.

# Acknowledgements

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