WACCM-X Tutorial Session

- Last CEDAR tutorial in 2018: CESM2.0/WACCM-X with CAM4/5 physics.
- New dynamical cores, physics packages (CAM6 and CAM7), and updated chemistry.
- On the verge of CESM3 release.
- Now available at NASA CCMC.

Agenda:

- What is the NSF NCAR Community Earth System Model and WACCM-X? What are the major model capabilities? (*Nick Pedatella*)
- How to download, set up, build and run the model? (Francis Vitt & Joe McInerney)
- What can NASA CCMC do for you? (Jack Wang)
- How to request runs at NASA CCMC? (Jack Wang).

Google meet link: meet.google.com/atz-qqmp-gwh

Office Hour: 4-5pm today in the multi-function room.





Nick Pedatella, Hanli Liu, Joe McInerney, and Francis Vitt High Altitude Observatory, NSF NCAR

> 25 June 2025 CEDAR Workshop



Outline

- Overview of CESM
- WACCM-X overview and capabilities
- Recent/upcoming developments

WACCM-X Overview & Science Goals

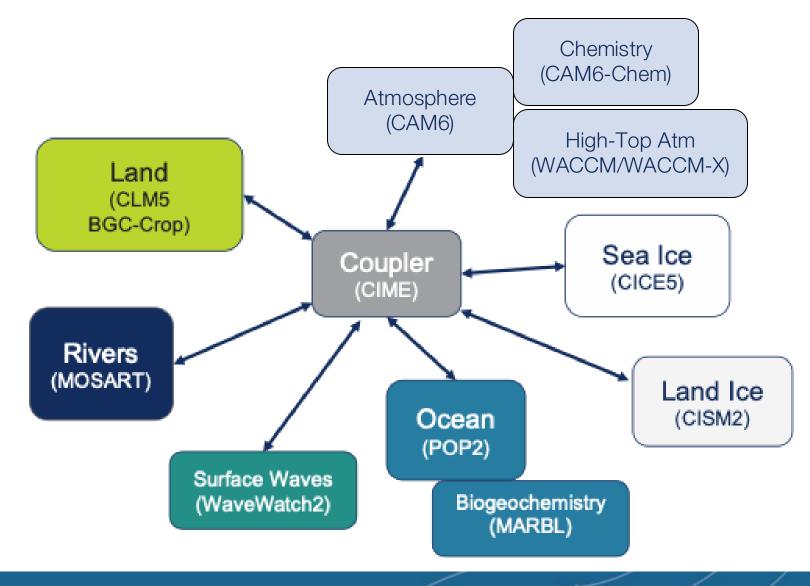
- WACCM-X is one of the atmospheric components of the NSF NCAR Community Earth System Model (CESM)
- Models the chemistry, dynamics, and physics of the atmosphere-ionosphere from the surface to the upper thermosphere
- Science goals:
 - Impacts of the Sun on the Earth system
 - Understand and quantify couplings between atmospheric layers through chemical, physical, and dynamical processes
 - Implications of the coupling to climate (downward coupling) and to the space environment (upward coupling)

WACCM-X is one of the atmosphere components of the NSF NCAR Community Earth System Model (CESM)

External Forcings:

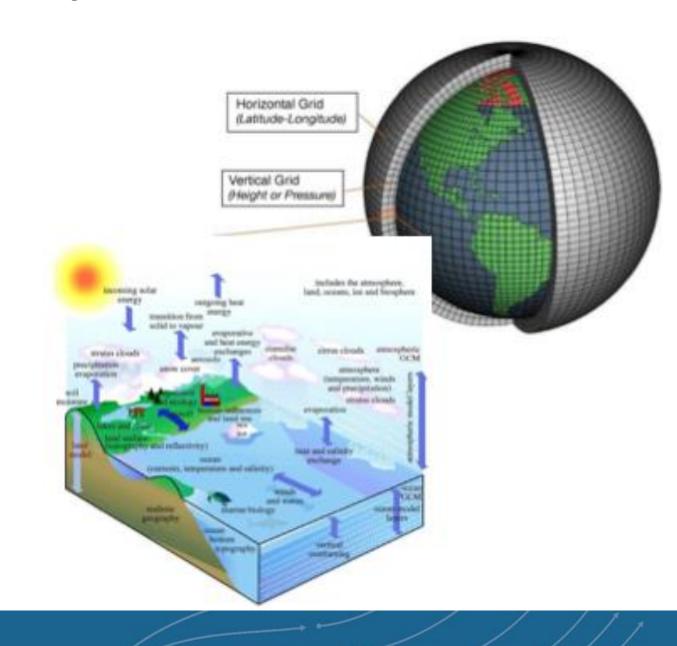
- Greenhouse gases
- Anthropogenic aerosols
- Volcanic eruptions
- Solar variability

. . .



What is an Earth System model?

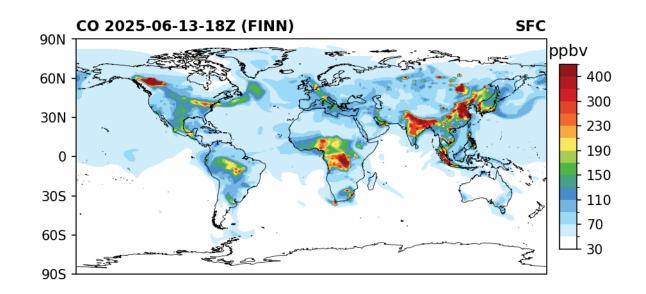
- Use physical equations to simulate key fields and processes in the atmosphere, ocean, land, sea-ice, land-ice, etc. and their couplings
- Processes that are below the grid resolution need to be parameterized
- Build on our understanding of processes from observations and highly-detailed models

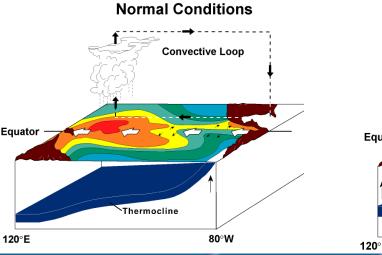


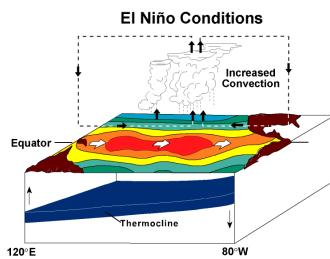
Applications of Earth System Models

- Earth system variability and change
- Air quality
- Biogeochemical cycles
- Sun-climate connections and space weather
- Process understanding
- Land-atmosphere and ocean-atmosphere interactions
- Ice sheet climate interactions
- Hydrology and ecology
- Earth system predictability
- Weather from local to global scales

• ...











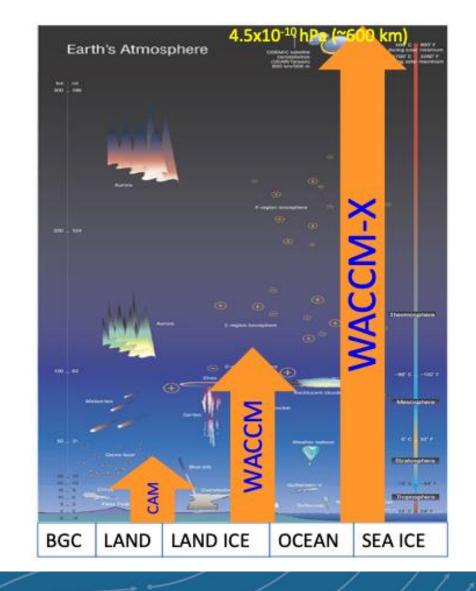
CESM has a large community of developers and users, including large simulation datasets and support for community users

- CESM is based on over 30 years of model development
- A true community model with external developers and support for users
- Community participation through working groups (2000+ researchers)
- Widely used, e.g., 400+ talks using CESM at AGU 2024
- 2.8 million lines of Fortran code, > 1000 geophysical variables
- Available on github: https://github.com/NCAR/cesm



WACCM-X is built upon the underlying CESM infrastructure and the CESM atmospheric models

- WACCM-X leverages the capabilities and developments in the lower-middle atmosphere that are part of the Community Atmosphere Model (CAM) and Whole Atmosphere Community Climate Model (WACCM)
- Detailed treatment of troposphere, stratosphere, and mesosphere processes, including detailed chemistry, are included in WACCM-X
- Dynamical core and physics developments and improvements can be adopted by WACCM-X
- Designed for climate, not necessarily weather timescales



Outline

- Overview of CESM
- WACCM-X overview and capabilities
- Recent/upcoming developments

Major CESM WACCM/WACCM-X Components

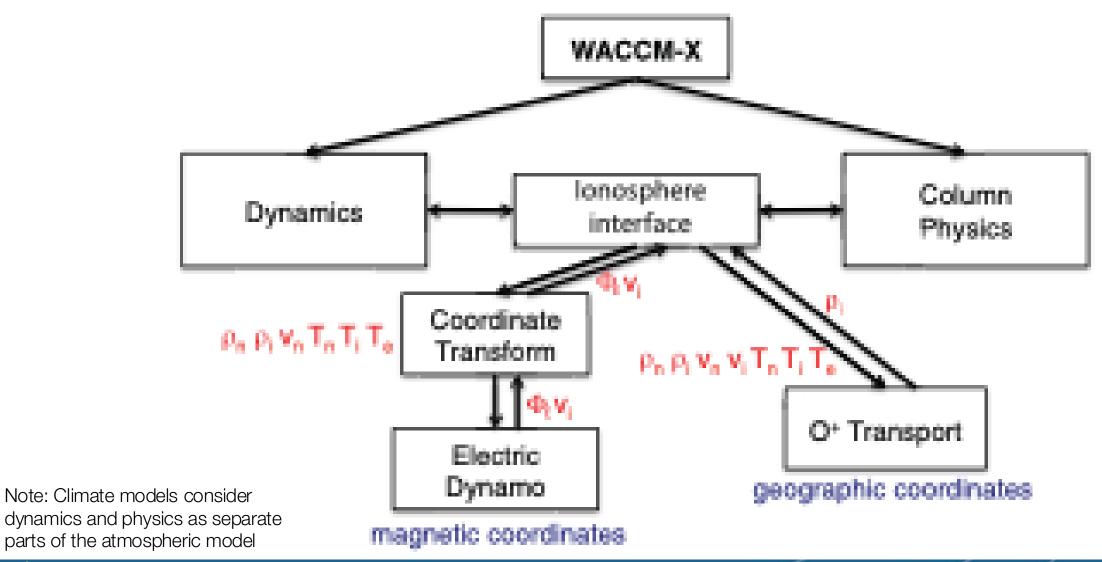
Model Framework	Chemistry	Neutral Atm. Physics	Ionosphere Physics	Resolution
Atmosphere component of NCAR CESM Extension of the NCAR Community Atmosphere Model (CAM) Finite Volume Dynamical Core (modified to consider species dependent Cp, R, m)	MAVIon Chemistry (~100 species) Fully-interactive with dynamics	Long wave/short wave/EUV RRTMG IR cooling (LTE/non-LTE) Modal Aerosol Model (MAM) CARMA Convection, precip., and cloud parameterization (CLUBB) Parameterized GW Major/minor species diffusion (+UBC) Horizontal/Vertical molecular viscosity and thermal conductivity (+UBC) Species dependent Cp, R, and m	Parameterized electric field at high latitudes. IGRF geomagnetic field Auroral processes, ion drag and Joule heating lon/electron energy equations Ambipolar diffusion lonospheric dynamo Coupling with plasmasphere/magnetosphere	Horizontal (lat x lon): 1.9 x 2.5 0.9 x 1.25 0.47 x 0.625 Vertical: 130, 145 levels (0 - ~600 km)



Key WACCM-X Capabilities

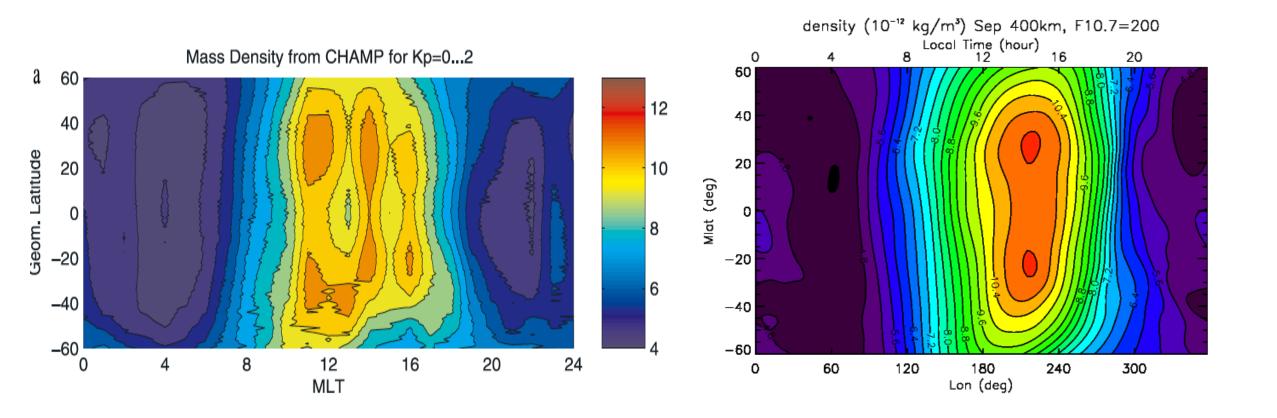
- Fully interactive chemistry, including ion chemistry
- Ionospheric electrodynamics using fully interactive dynamo
- Ion transport in the F-region
- Solar inputs from F107/EUVAC or FISM for solar flares
- Magnetospheric inputs using empirical (Heelis/Weimer) or specifications (AMIE, GAMERA)
- Coupling with SAMI3 plasmasphere model (NRL/Syntek Tech.)
- Specified dynamics to constrain meteorology up to ~50 km
- Whole atmosphere data assimilation for specification and forecast using DART (Data Assimilation Research Testbed)

WACCM-X integrates the model dynamics and physics through an ionospheric module that enables electrodynamics and O+ transport

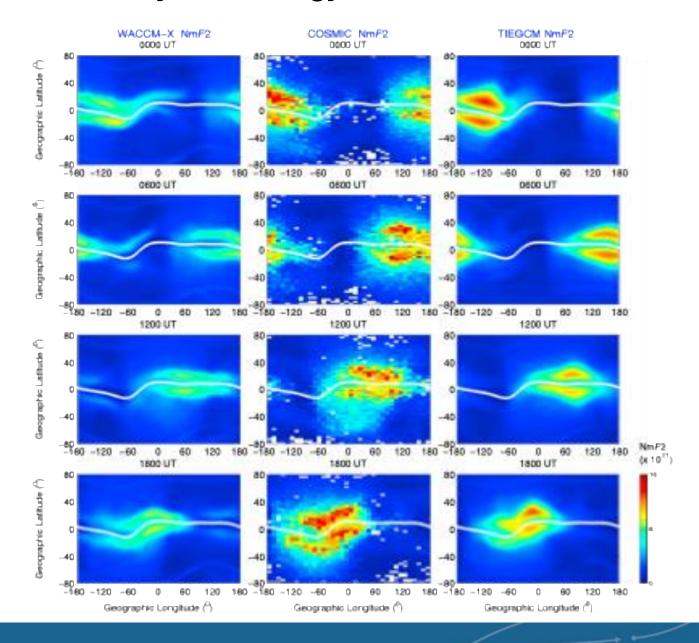




WACCM-X simulated thermosphere mass density is in general agreement with CHAMP satellite observations



Morphology of electron density climatology is consistent with observations and TIEGCM

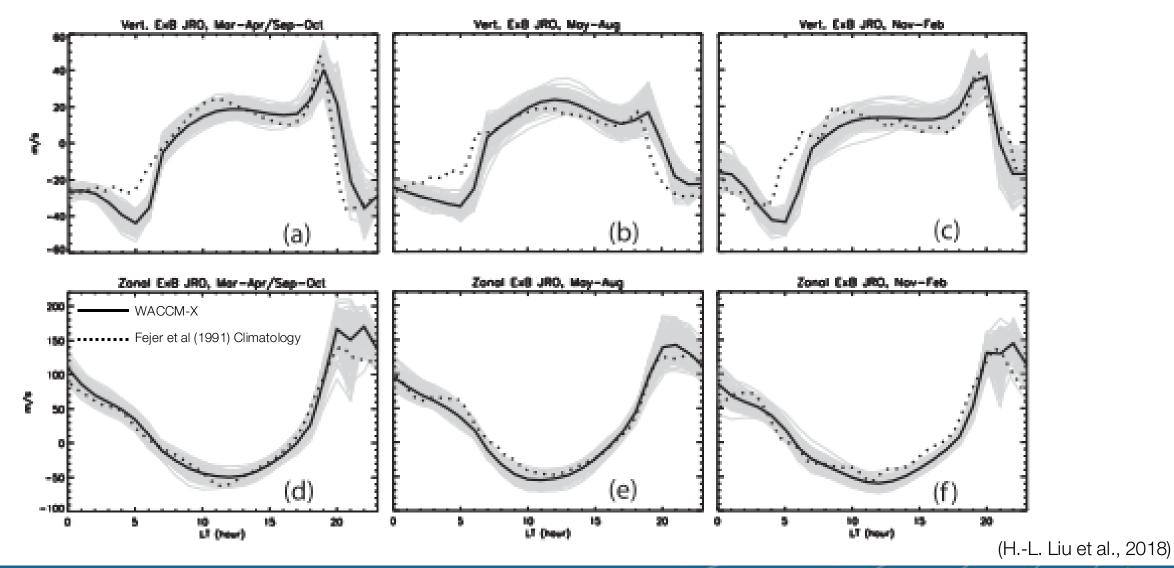


*Underestimation of NmF2 due to excessive GW diffusion WACCM-X v2.0 is largely fixed in v2.1

(J. Liu et al., 2018)



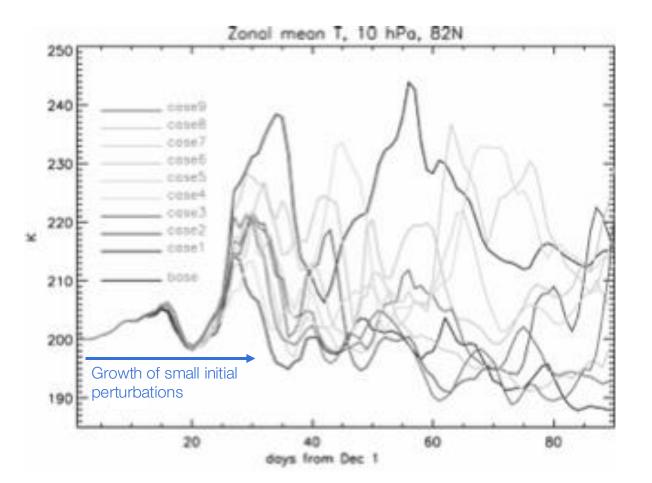
Simulated equatorial vertical drifts in WACCM-X generally agree with climatology and exhibit day-to-day variability due to lower atmosphere variability





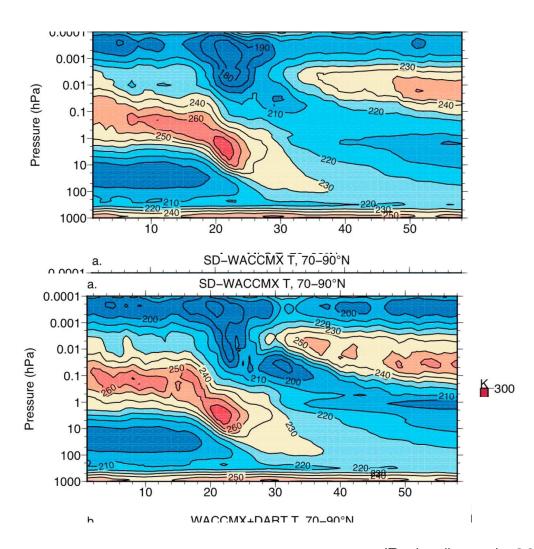
Lower atmosphere: Free-run versus Specified Dynamics

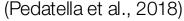
- WACCM-X can be configured as either a freerun or with specified dynamics
- Free-run simulations are unconstrained, and the model generates its own "weather" that is representative of climatological behavior
- Free-running simulations are useful for studying general behavior of the system, longterm simulations and/or time periods when reanalyses are unavailable
- Because of internal model variability, ensemble simulations are often required



Lower atmosphere: Free-run versus Specified Dynamics

- WACCM-X can be configured as either a freerun or with specified dynamics
- Specified Dynamics (SD) simulations constrain the model meteorology to external fields. $X_{model} = \alpha(X_{met} X_{model})$
- Default SD configuration is to nudge the model fields to MERRA-2 reanalysis up to ~50 km
- Assumption is that middle-upper atmosphere variability is controlled by dynamics in the troposphere-stratosphere
- Useful for simulating specific time periods and/or comparing to observations



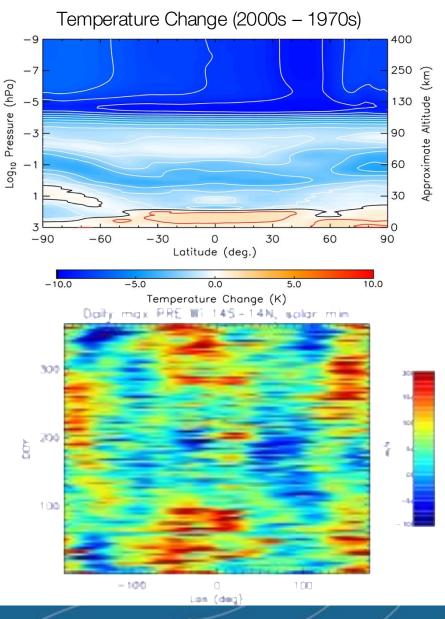




Example science use cases of WACCM-X

- Day-to-day variability in the ionosphere and thermosphere
- Long-term trends (historical and projection)
- Sudden stratospheric warmings: dynamics, chemistry, and ionospheric effects
- Effects of gravity waves, tides, and planetary waves on the circulation and chemistry of the MLT
- Geomagnetic storms

• ...



(Liu et al., 2018; Solomon et al., 2018)



Outline

- Overview of CESM
- WACCM-X overview and capabilities
- Recent/upcoming developments

From CESM2 to CESM3

- CESM2 was released in 2018 and included the initial release of WACCM-X
- Subsequent CESM2.x model releases include developmental (i.e., not scientifically supported)
 updates and bug fixes
- NSF NCAR intends to release CESM3 in the near future
- CESM3 will include many new features, including a new dynamical core and significant updates to the atmospheric physics and chemistry (CAM7/WACCM7/WACCM7-X)

Major CESM WACCM/WACCM-X Components

Model Framework	Chemistry	Neutral Atm. Physics	lonosphere Physics	Resolution
Atmosphere component of	MA/Ion Chemistry (~100	Long wave/short wave/EUV	Parameterized electric field at	Horizontal (lat x lon):
NCAR CESM	species)		high latitudes.	1.9 x 2.5
		RRTMG	1005	0.9 x 1.25
Extension of the NCAR	Fully-interactive with		IGRF geomagnetic field	0.47 x 0.625
Community Atmosphere	dynamics	IR cooling (LTE/non-LTE)	Auroral processos ion drag	NE16, 30, 60, 120
Model (CAM)		Modal Aerosol Model (MAM)	Auroral processes, ion drag and Joule heating	NETO, 30, 60, 120
Finite Volume Dynamical		Modal Aerosol Model (MAM)	and Joule nealing	120 km
Core (modified to consider		CARMA	lon/electron energy	120 KIII
species dependent Cp, R,			equations	Vertical:
m)		Convection, precip., and cloud	oquationis	70, 135 levels
,		parameterization (CLUBB)	Ambipolar diffusion	(0 – 140 km)
Species dependent spectral		,	·	,
element (SE) dynamical		Parameterized GW (+ MM	lonospheric dynamo	130, 145 levels
core		parameterization)		(0 - ~600 km)
			Coupling with	
Regridding scheme		Major/minor species diffusion	plasmasphere/magnetospher	130, 189, 273
A ADA G. A		(+UBC)	е	(0 - ~600 km)
MPAS-A				
		Horizontal/Vertical molecular		
CESM2 WACCM-X		viscosity and thermoal conductivity (+UBC)		
CESM3 WACCM-X				
WACCM Development		Species dependent Cp, R, and m		

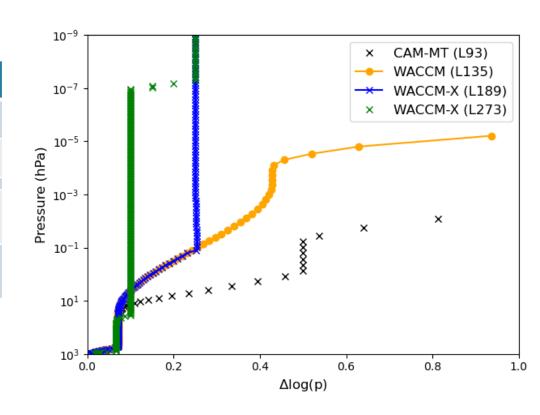


WACCM & WACCM-X for CESM3

	Dyn. Core	Resolution	# levels	Chemistry	Physics
WACCM6	FV	1°, 2°	70 (110)	TSMLT, MA	CAM6
WACCM7	SE	ne30, ne16	135	t1ma, t4ma	CAM7
WACCM-X 2.1	FV	1°, 2°	130	TSMLT, MA, IT chemistry	CAM6
WACCM-X v?	SE	ne30, ne16 (ne120)	189 (273)	t1ma, t4ma IT chemistry	CAM7



- Good QBO at both 1° and 2° resolutions
- Seasonal cycle of winds/temperatures in the stratospheremesosphere and the impact on chemistry
- Thermosphere extension of SE dycore, ionosphere thermosphere climatology



Extension of Spectral Element (SE) dynamical core into the thermosphere

- Spectral Element (SE) dynamical core will be used by future versions of CESM atmosphere models (CAM, WACCM, WACCM-X)
- Cubed sphere grid that eliminates polar singularity
- Capability for high-resolution and nested grids
- SE dynamical core has been extended into the thermosphere: species dependence; horizontal molecular viscosity/diffusion



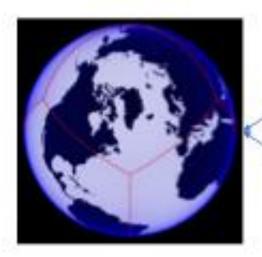


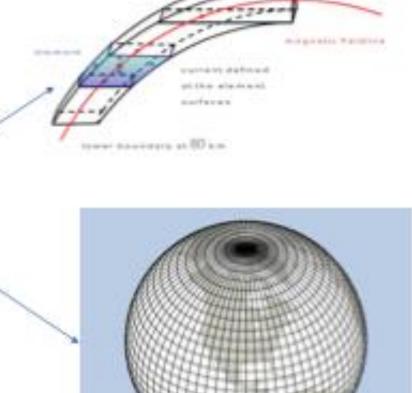
New regridding scheme for mapping between regular/irregular grids

 Earth System Modeling Framework (ESMF) based regridding scheme maps fields between physics mesh and regular/irregular grids

Independent of dynamical core grid





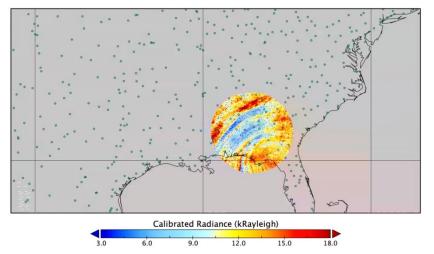


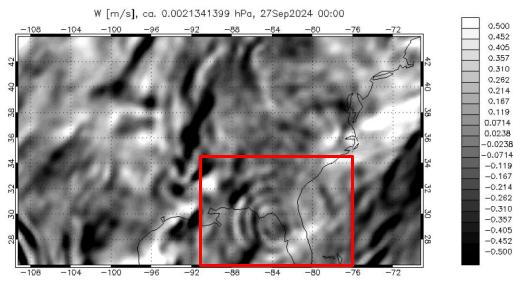
Implementation of meteorological constraint for high-resolution (ne120, ~0.25°) WACCM-X

- Recently implemented physics-side nudging for high-resolution WACCM-X simulations
- Enables simulations of the day-to-day variability of small-scale waves, and their impacts, on the middle and upper atmosphere
- Capability enables new scientific studies and is used to support current and future satellite missions

Gravity Waves from Hurricane Helene

NASA AWE Observation 02:49 UT, Sept 27, 2024 (~87 km)





Credit: Hanli Liu (HAO), Jiarong Zhang (USU)



Chemistry Updates for the Mesosphere and Thermosphere

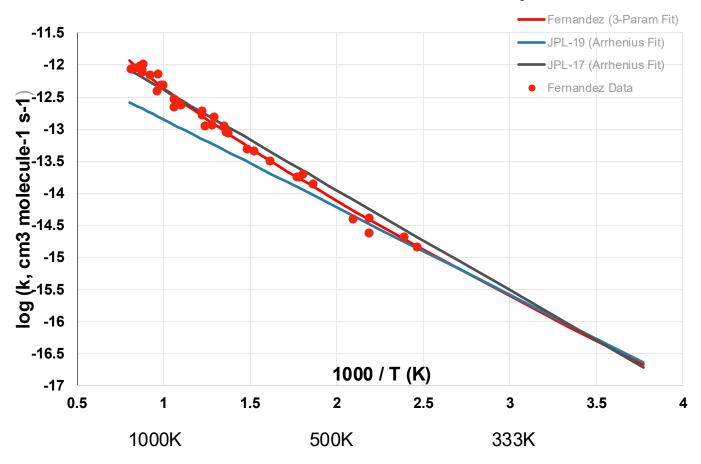
	Reaction	Original Rate Constant (A, Ea/R)	Updated	Comment
Updated rates	$N(^4S) + O_2 -> NO + O$	JPL-19 3.3e-12, 3150	3-parameter fit, Next slide	Fernandez et al., 1998
	$N(^2D) + O_2 -> NO + O(^1D)$	5e-12	K=6.2e-12 (T/300.)	Duff et al., 2003
New reactions	$N(^{2}D) + NO \rightarrow N_{2} + O(^{1}D)$	NA	K = 7.3e-11	Roble, 1995
	$N(^2D) -> N(^4S) + hv$	NA	K = 1.06e-5	Roble, 1995
	$N(^{2}D) + e \rightarrow N(^{4}S) + e + 2.38 eV$	NA	$K = 3.6e-10_r8 * (Te/300.0_r8)^{1/2}$	Roble, 1995

Credit: Doug Kinnison, Jun Zhang (ACOM)



Updated N + O₂ Reaction Rate

N + O2 Kinetics Measurement Summary

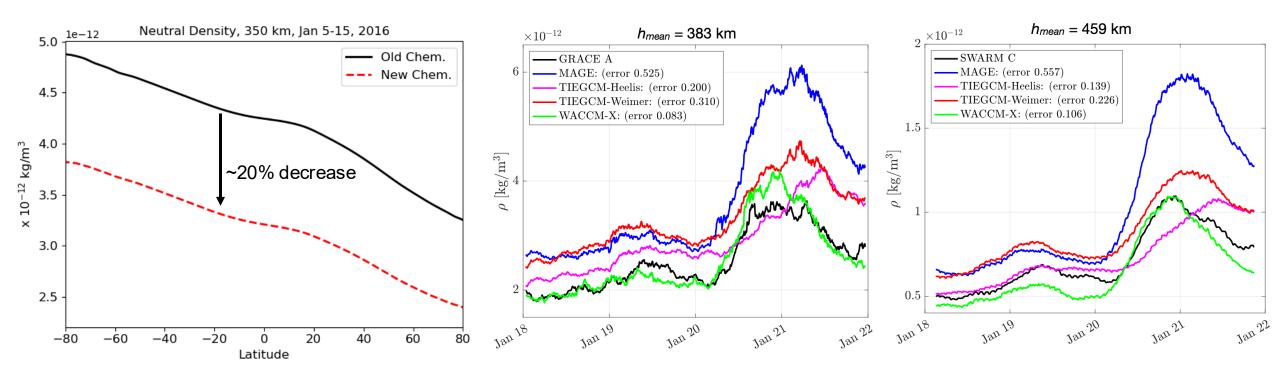


- Decrease in reaction rate from JPL-17 (CESM2) to JPL-19 leads to less NO production at high temperatures
- NO cools the thermosphere, so less NO leads to larger temperatures and neutral densities

Credit: Doug Kinnison, Jun Zhang (ACOM)



WACCM-X simulated neutral densities with updated reaction rates show good agreement with observations

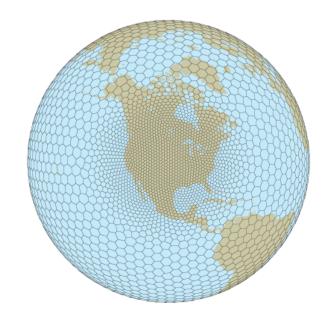


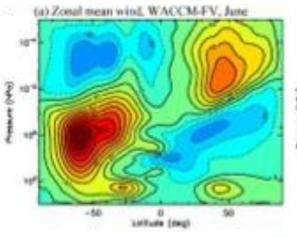
Credit: Jordi Vila Perez (HAO)

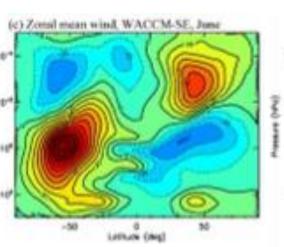


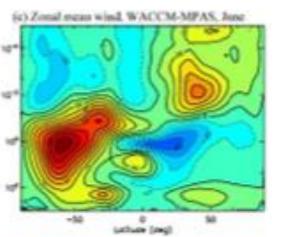
Model for Prediction Across Scales (MPAS) dynamical core

- MPAS-A is a non-hydrostatic dynamical core solved on centroidal Voronoi mesh
- MPAS has been extended into the lower thermosphere (Kamali et al., 2024)
- Further developing the MPAS dynamical core for future use as the dynamical core in WACCM-X





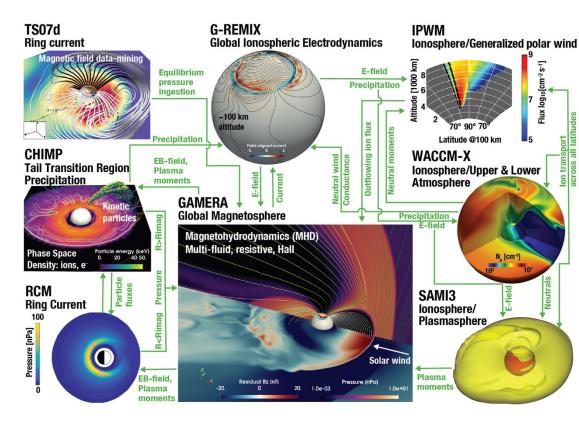




WACCM-X/GAMERA: Towards a whole geospace model

NASA Center for Geospace Storms (CGS) and HAO have recently completed *two-way coupling* between WACCM-X and the GAMERA magnetosphere model, paving the way for a whole geospace model

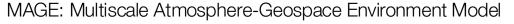
Model Framework for CGS MAGE













Summary

- CESM/WACCM-X is an open source, community, model of the whole atmosphere-ionosphere coupled system
- CESM2 (WACCM-X v2.x) has significantly advanced scientific understanding of day-to-day variability, SSWs, wave-driven variability, long-term trends, ...
- Release of CESM3 (WACCM7-X) will provide new modeling capabilities, including SE dynamical core that enables high-resolution simulations
- CESM is a community model and we welcome scientific and development collaborations

References



WACCM-X Tutorial

CEDAR 2025

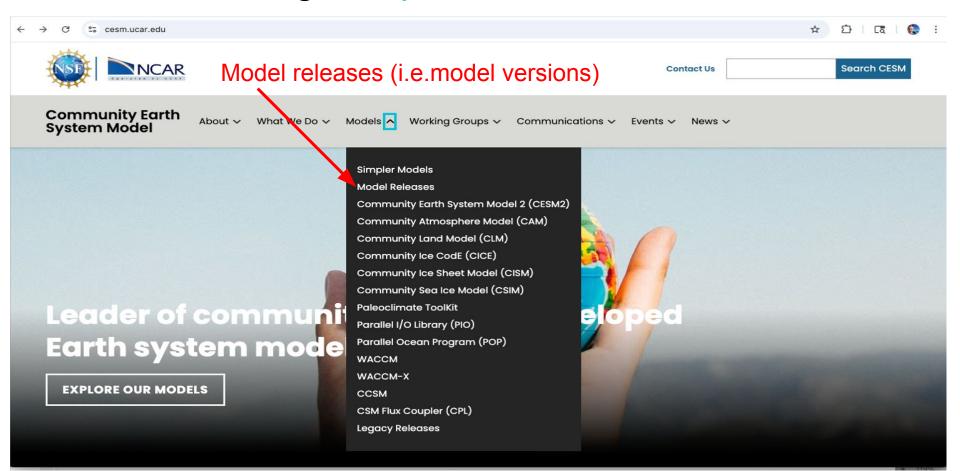


Overview

- Downloading model source code
- Creating a case directory
- Case setup, build, and run
- Contents of directories
- Modifying case settings
- Description, access, and viewing model output
- Documentation and help

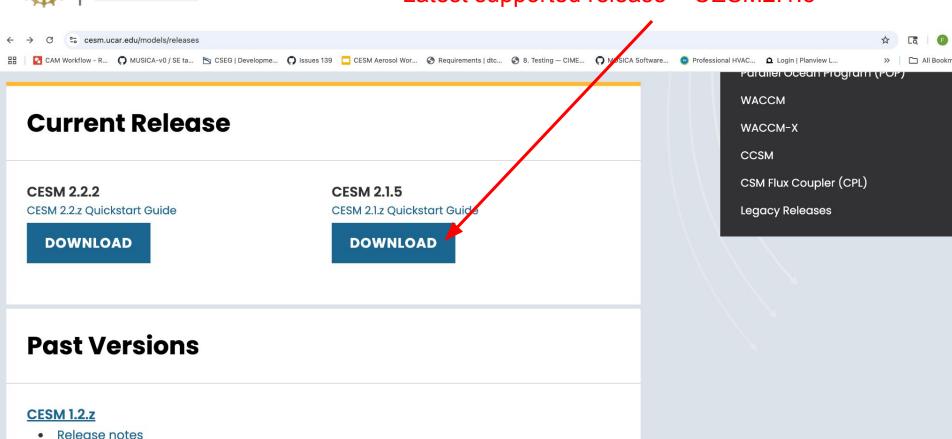


CESM Web Page: http://www.cesm.ucar.edu



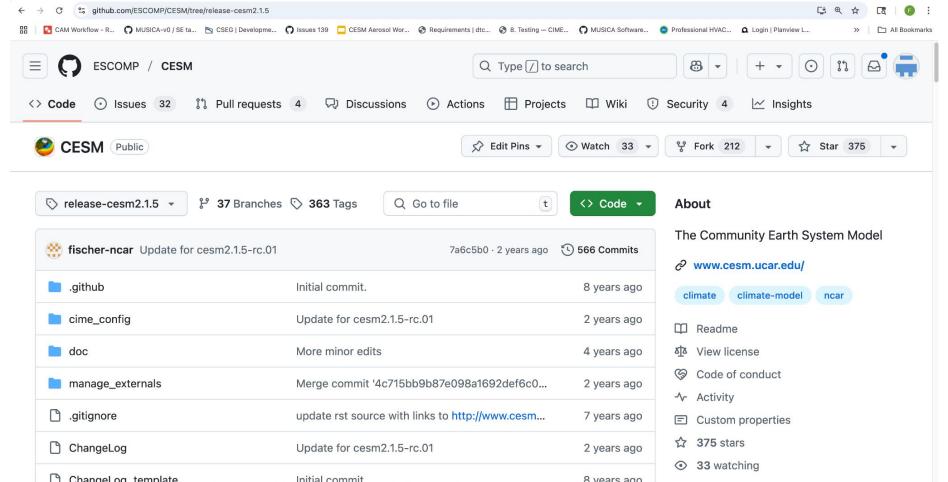


Latest supported release – CESM2.1.5



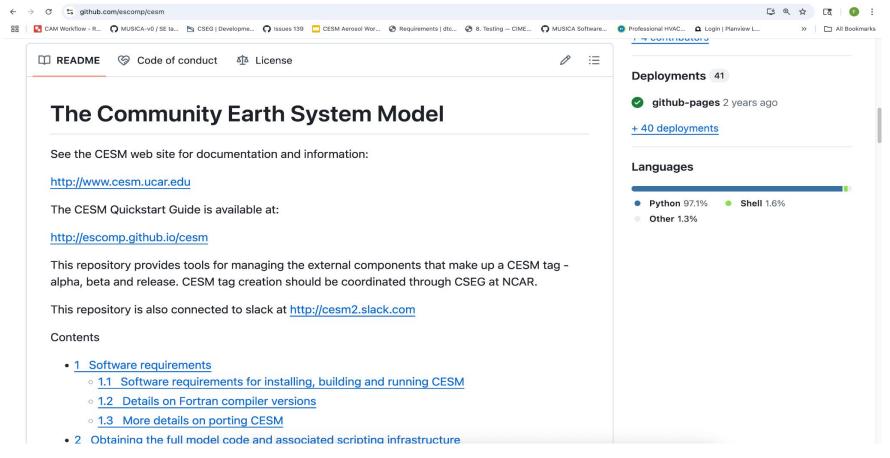
CESM github repository





See the README





Download and Basics

Download from github

```
> git clone https://github.com/ESCOMP/CESM.git -b cesm2.1.5 cesm2.1.5
> cd cesm2.1.5
> manage_externals/checkout_externals
```

Not covered here for unsupported machines (.e.g. derecho)

- Porting
- Create input data root directory

Basic steps for running

- Create case
- Invoke case.setup, case.build
- Submit the run ...

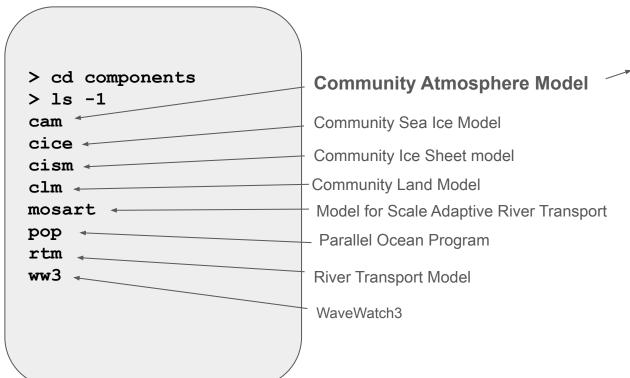


Download listing of CESM

```
> cd cesm2.1.5
> 1s -1
                                           Scripts for creating case, etc.
ChangeLog
                                           under cime
ChangeLog template
cime
cime_config
components
describe version
doc
                                           Model source code under
Externals.cfg
                                           components
LICENSE. txt
manage externals
README.rst
```



Components listing



> cd cam/src > ls -1 advection chemistry control cpl dynamics ionosphere physics unit_drivers utils

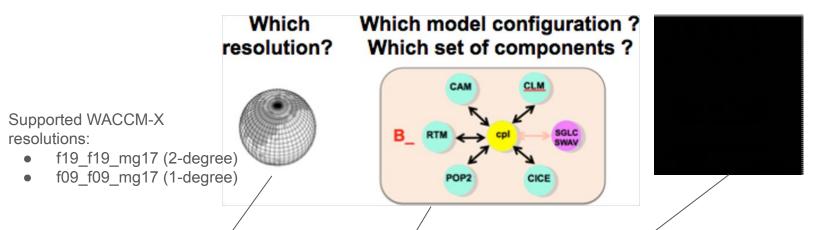


Create a new case



In the **cime/scripts** directory, **create_newcase** is the tool that generates a new model case.

create_newcase requires 3 arguments:



- > ./create newcase --res f19 f19 mg17 --compset FXHIST --case /glade/work/fvitt/cesm/cases/fe21.FXHIST.f19.t001
- > ./create_newcase --help

What is a compset?



Sea Ice

Waves

(WW3)

Ocean

(POP2)

Ecosystem (MARBL)

Atmosphere

MCT Driver/ Mediator

Land Ice

(CISM)

River Runof

(MOSART)

Land (CLM5)

"FXHIST" is an example of a component set, or "compset", which defines the configuration of the CESM component models: atmosphere, land, ocean, sea ice, and land ice.

All WACCM-X components use non-interactive data models for ocean and sea ice, and do not include interactive land ice. Such compsets all begin with the letter "F".

To list available WACCM-X compsets, under cime/scripts enter:

> ./query_config --compsets | grep %WXIE

Short name

Long name

: 2000 cam40%wxie_clm40%sp_cice%pres_docn%dom_rtm_sglc_swav

FXHIST : HIST_Cam40%wxie_clm40%sp_cice%pres_docn%dom_rtm_sglc_swav

FXmadHIST : HIST_Cam40%wxied_clm40%sp_cice%pres_docn%dom_rtm_sglc_swav

FXSD : HIST_Cam40%wxie%sdyn_clm40%sp_cice%pres_docn%dom_rtm_sglc_swav

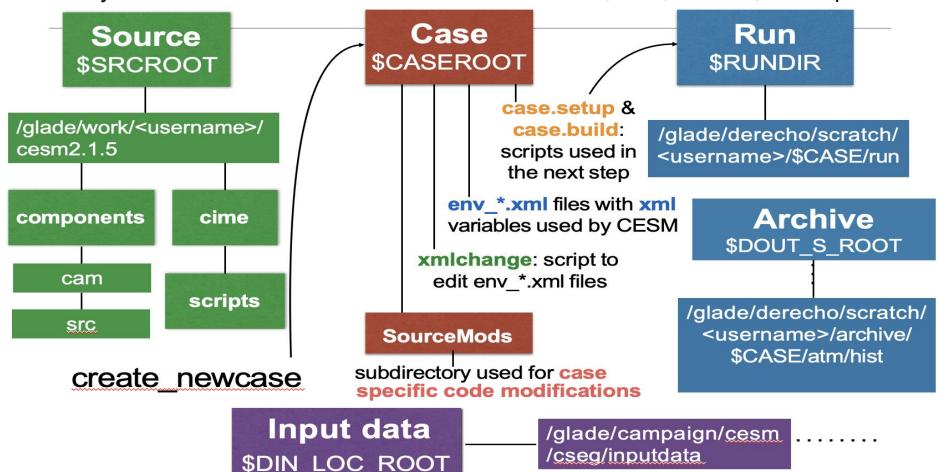
FXmadSD : HIST_Cam40%wxied%sdyn_clm40%sp_cice%pres_docn%dom_rtm_sglc_swav

WACCM-X Compsets



FX2000	Perpetual year 2000 emissions forcings	Middle atmosphere WACCM chemistry
FXHIST	Transient emissions forcings	Middle atmosphere WACCM chemistry
FXmadHIST	Transient emissions forcings	Middle atmosphere chemistry plus D-region ion chemistry
FXSD	Transient emissions forcings Nudged to MERRA meteorology	Middle atmosphere WACCM chemistry
FXmadSD	Transient emissions forcings Nudged to MERRA meteorology	Middle atmosphere chemistry plus D-region ion chemistry

Directory Structure of Model Source and Created Case, Run, Archive, and Input Data



Compiling: Setup and Build (Derecho)

After creating your case, go to the case directory:

```
NSF
```

```
> cd /glade/work/fvitt/cesm/cases/fe21.FXHIST.f19.t001
```

Setup the case:

> ./case.setup

Build the case:

>./case.build

Problems? Try:

- >./case.setup --reset
- >./case.build --clean-all
- >./case.build

Running: Submit and Checking Output (Derecho)

After compiling and building the case, submit the job:

```
> ./case.submit
```

Check the run in the queue:

```
> qstat -u fvitt
```

Check run directory while running::

```
> ls /glade/derecho/scratch/fvitt/fe21.FXHIST.f19.t001/run
```

Check archive directory after run completes::

```
> ls /glade/derecho/scratch/fvitt/archive/fe21.FXHIST.f19.t001/atm/hist
```

Problems? Log files in run directory can help: cesm*log*, atm*log*, cpl*log*



After running out-of-the-box, can make changes to run configuration

XML Files (build and run control variables)

There are multiple env_*.xml files in the \$CASEROOT directory:

- env_archive.xml: specifies rules for short term archive script case.st_archive
- env_batch.xml: specifies batch specific settings used in case.submit script
- env_build.xml: specifies build information used in the case.build script
- env_case.xml: set by create_newcase and cannot be modified
- env_mach_pes.xml: specifies PE layout on NCAR HPC for components and used by case.run script
- env_mach_specific.xml: specifies machine specific information used in case.build script
- env_run.xml: sets run time information (such as length of run, number of submissions, ...)

Use xmlquery tool to see settings

```
>./xmlquery --partial STOP --full
```

Use xmlchange to change settings:

>./xmlchange \$CONTINUE_RUN="TRUE"



Component processor usage (or layout)

```
> ./pelayout
     NTASKS
             NTHRDS
                     ROOTPE PSTRIDE
Comp
       1536/
CPL:
                 1;
       1536/
                 1;
ATM:
       1536/ 1;
LND:
ICE :
       1536/
OCN:
       1536/
ROF :
       1536/
       1536/
                 1;
GLC :
WAV:
       1536/
       1536/
ESP:
ESMF AWARE THREADING is False
ROOTPE is with respect to 128.0 tasks per node
```

To change number of MPI tasks:

```
> ./xmlchange NTASKS=-10
```

Negative value indicates number of compute nodes



Modifying Model Settings In user_nl_cam File

CaseDocs/atm_in file has current settings which can be modified in user_nl_cam

```
> less $CASEROOT/CaseDocs/atm_in
```

Output fields:

Search for "MASTER FIELD LIST" in atm log file output in archive directory

```
> less /glade/derecho/scratch/fvitt/archive/fe21.FXHIST.f19.t001/logs/atm*log*
```

Modified by setting "fincl" namelist variables

```
fincl2= 'Z3', 'T', 'TIon', 'TElec', 'e', 'U', 'V', 'OMEGA', 'UI', 'VI',
'WI', 'EDens', 'ElecColDens', 'PHIM2D', 'PS', 'EDYN_ZIGM11_PED',
'EDYN_ZIGM2_HAL', 'ED1', 'ED2', 'O', 'O2', 'H', 'NO',
'TTGW', 'UTGW_TOTAL', 'DTCOND', 'QRS', 'QRL', 'QNO'
```

Output frequency, times per file, averaged or instantaneous

```
nhtfrq = 0, -3 , -24, -24, -120, -24
mfilt = 1, 8, 7, 7, 10, 365
avgflag_pertape= 'A', 'I', 'I', 'A', 'A', 'A'
```



Characteristics of WACCM-X Output netCDF History Files

- netCDF: self-describing binary data format used for primary CESM output
- History files: WACCM-X output is written to several output streams, each
 with a particular frequency and averaging characteristic which can be modified
 in the user_nl_cam file
 - h0: monthly averages
 - f.e22.FXSD.f09_f09_mg17.001.cam.h0.2000-01.nc (January 2000)
 - f.e22.FXSD.f09_f09_mg17.001.cam.h0.2000-02.nc (February 2000)
 - h1: 3-hourly instantaneous
 - f.e22.FXSD.f09_f09_mg17.001.cam.h1.2000-01-01-00000.nc (January 1, 2000)
 - f.e22.FXSD.f09_f09_mg17.001.cam.h1.2000-01-02-00000.nc (January 2, 2000)
 - h2: daily instantaneous
 - h3: daily averages
 - h4: 10-day average tidal coefficients
 - h6: daily averages, zonal mean circulation diagnostics



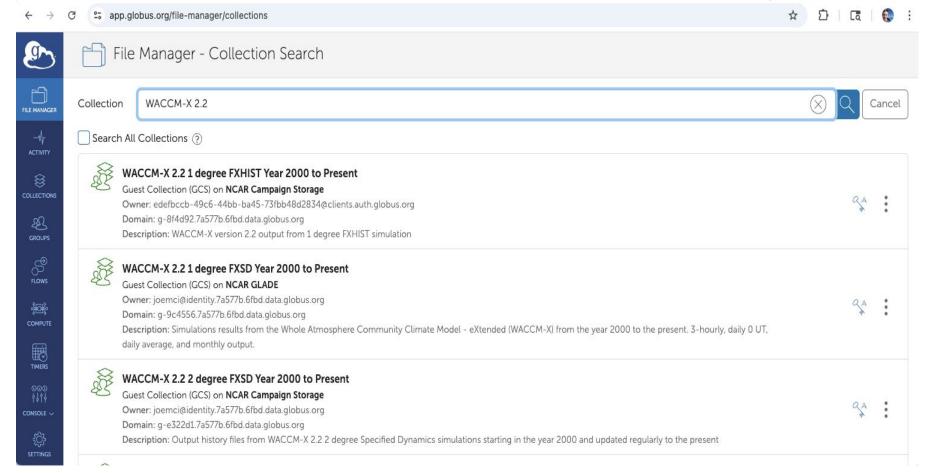
WACCM-X Output on Globus (1)

- WACCM-X output is available Globus website globus.org
- Three collections of WACCM-X output from 2000 through 2024
- 1 degree horizontal resolution historical (FXHIST) simulation output
- 1 (f09) and 2 (f19) degree horizontal resolution specified dynamics (FXSD) simulation output
- Updated regularly
- If new to Globus, go to https://www.globusid.org/create to sign up
- Log in, choose "File Manager" tab on left then in "Collection", search for "WACCM-X 2.2" (next slide)

WACCM-X Output on Globus (2)







University Allocations

<u>https://arc.ucar.edu/docs</u> - Left side bar "Getting Started" > "Allocations" > University Allocations"

The table below summarizes the HPC resource limits for each type of allocation. For Small, Exploratory, and Classroom projects, the amounts shown assume use of only Derecho or Derecho GPU. For requests to use both parts of Derecho, proportionally smaller limits apply—e.g., up to half the Derecho limit and half the Derecho GPU limit can be requested together.

Allocation request	Initial HPC limit*	Supplement HPC limit	Frequency	Funding eligibility
Large	No upper limit (subject to availability)	No upper limit (subject to availability)	Semi-annual panel review	NSF award required
Small	Derecho: 1 million core-hours Derecho GPU: 2,500 GPU- hours	Derecho: 1 million core-hours Derecho GPU: 2,500 GPU- hours	Continuous	NSF award required
Exploratory & Classroom	Derecho: 500,000 core-hours Derecho GPU: 1,500 GPU- hours	Derecho: 500,000 core-hours Derecho GPU: 1,500 GPU- hours	Continuous	No external funding award
Data Analysis	n/a (Casper only)	n/a (Casper only)	Continuous	Any funding source

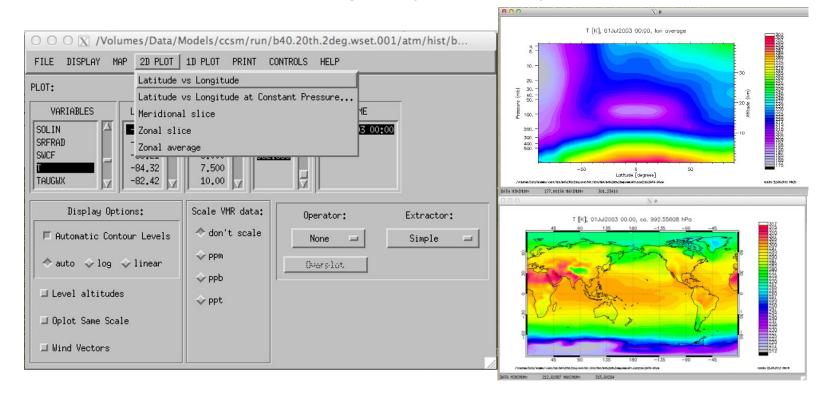
Submitting Your Request. For all types of university allocations, including any subsequent extension or supplement requests, requests should be submitted via the **ARC portal's Allocations section**. If you have questions about these options, please contact us via the **Research Computing help** desk.

Quick view of WACCM-X output





GEOV is an IDL-based viewer for geophysical history files



Quick view of WACCM-X output: GEOV



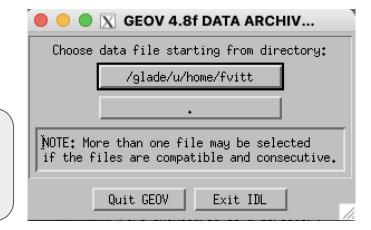
GEOV can be downloaded from github:

> git clone https://github.com/NCAR/GEOV.git

```
Can append GEOV path to idl_path:
  idl_path = expand_path('+/location/GEOV')
```

On NCAR machines (.e.g., casper or derecho)

> module load idl
> export IDL_STARTUP=~fvitt/idl_startup
> idl geov



Resources



Quick Start

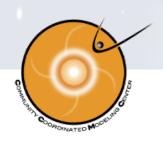
https://escomp.github.io/CESM/versions/master/html/index.html

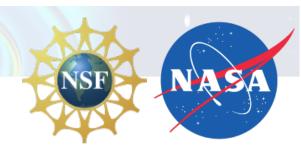
Online Tutorial

https://ncar.github.io/CESM-Tutorial/README.html

CESM Discuss Forums (get help)

https://bb.cqd.ucar.edu/cesm/





NASA Community Coordinated Modeling Center (CCMC) WACCM-X/RoR Overview

Jack Wang and all CCMC members

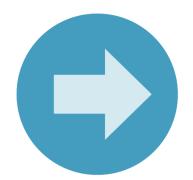
Community Coordinated Modeling Center, NASA GSFC

June 25nd, 2025

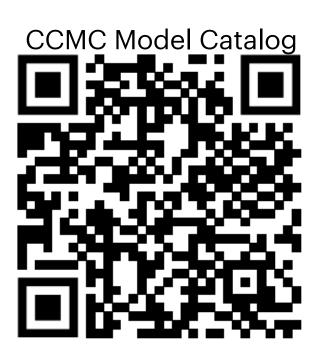
CCMC provides on-demand model simulations to support research studies, missions, and event analyses



Instant Run



Run-on-Request

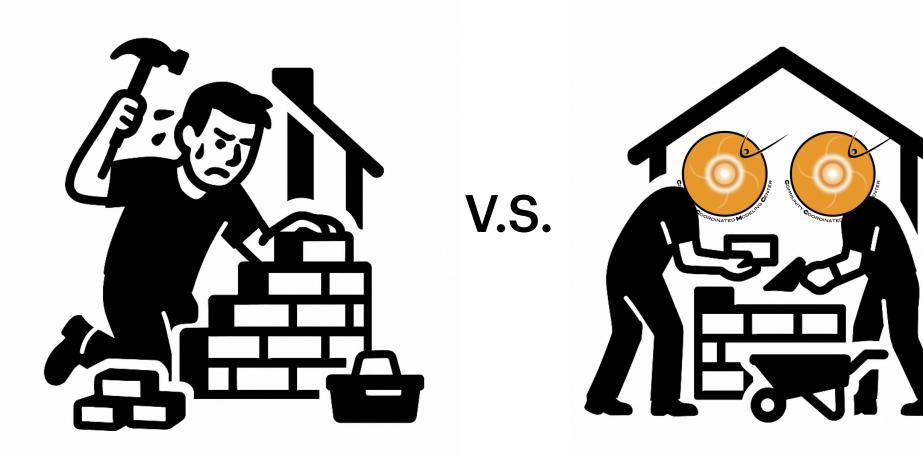




Continuous Run



CCMC builds bridges between model developers and end users



CCMC acknowledges all the model developers for the permission to use the models and software tools at CCMC





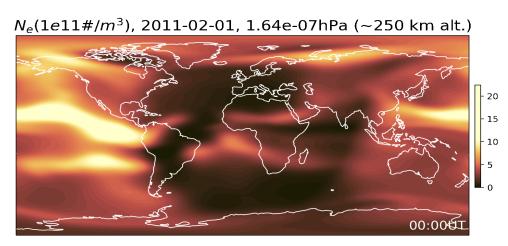
Runs-on-Request (RoR) - Execute simulations upon customer request

				•
		Plasmasphere	Cosgr	ove-PF
ANMHD		OpenGGCM	HWM14	GITM WAM-IPE
HeliosphericTomographyIPS CORHEL/MAS-TDM CORHEL/MAS		MAGE	DTM C1	TPe Weimer
		SWMF	JB2008	SAMI3 PBMOD
CORHEL-CME	DIPS	LFM	Ovation-Pri	
CORHEL-MAS_WSA		NAIRAS	MPTAM WACC	MX TIE-GCM
SWMF/AWSoM_R OSPREI PFSS NLFF WSA	EPREM	CalcDeltaB HYPE	RS-Global	
	GAMERA-Helio	VERB-3D R	AM-SCB	
	SEPMOD		PAMHD	
		GUMICS Tsygan	enko Magnetic Fie	ld
		VPIC_DIST		
Corona	Heliosphere	Magnetosphere	Inner	Ionosphere/

Magnetosphere Thermosphere

A Tutorial on the NSF NCAR WACCM-X: CCMC RoR Overview

WACCM-X (CESM2.2.0) is available through RoR since 2023, first whole atmosphere model at CCMC

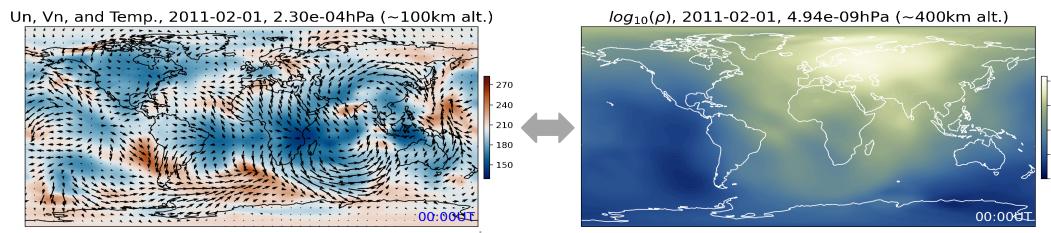


- output cadence: 10 mins.
- resolution: 0.9° Lat. x 1.25° Lon.
- variable: T, U, V, Z, drift, TEC, NE, composition...





ion-neutral coupling

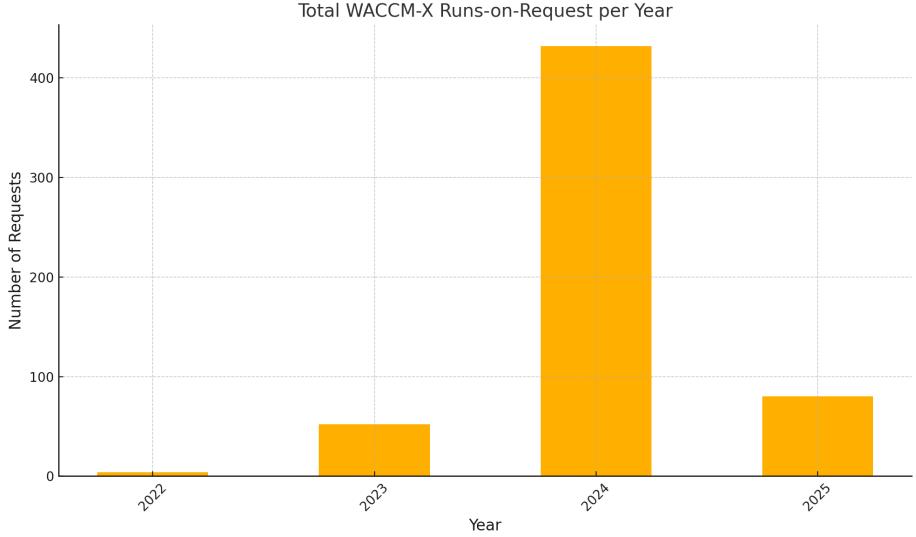


lower-upper atm. coupling





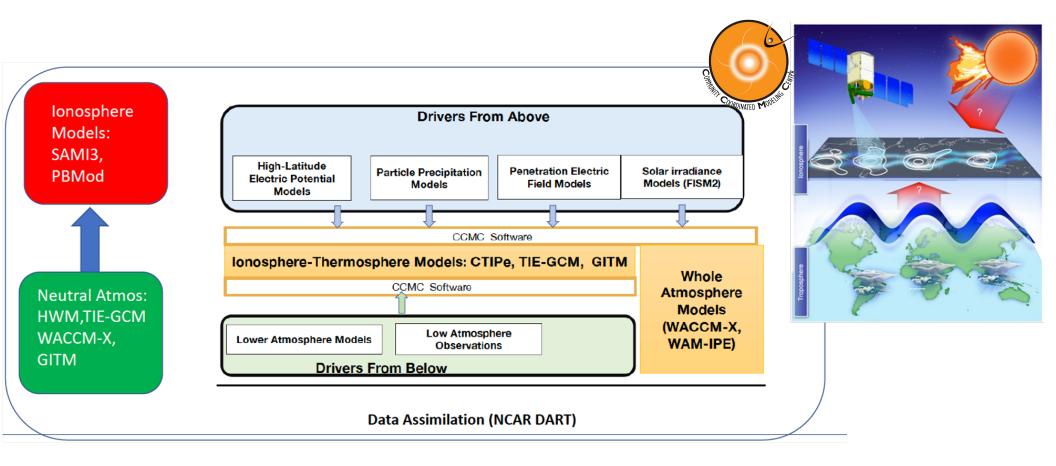
Over 560 WACCM-X requests since 2023 (from ~100 unique users across > 10 countries)







Models and services available at CCMC (thermosphere/ionosphere models coupling with magnetosphere MHD and lower atmosphere drivers) -- providing system solution

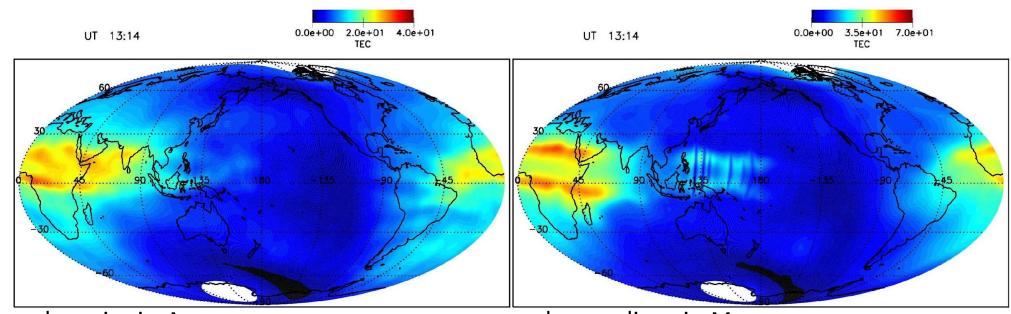






First global ionosphere model SAMI3 simulates weather of the ionosphere

- Options of empirical (HWM/MSIS) or physics-based model inputs (e.g., TIE-GCM, TIE-GCM/ICON, WACCM-X)
- SAMI3/WACCM-X is available through CCMC RoR.
 - enable to study day-to-day variability of plasma bubbles and TIDs

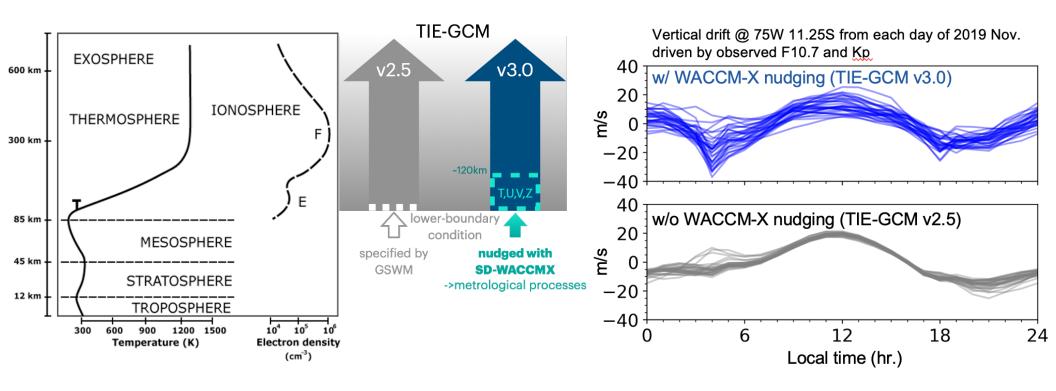


solar min. in Aug.

solar medium in Mar.



TIE-GCM v3.0 nudged with SD-WACCMX captures day-to-day variability in the lonosphere/Thermosphere

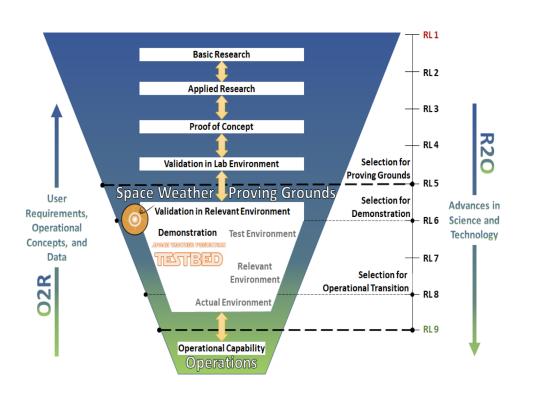


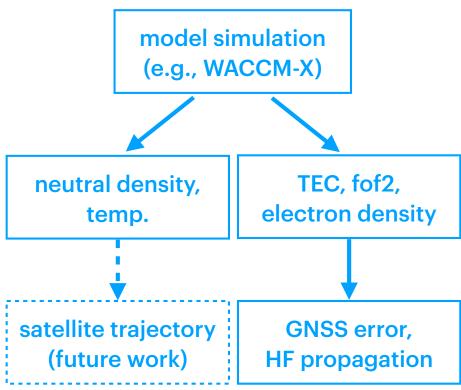
Enhancing the model capability to reproduce variability of the space environment as related to lower atmosphere forcing on day-to-day weather scales.





CCMC performs unbiased testing and validation of hosted models





impact-based validation



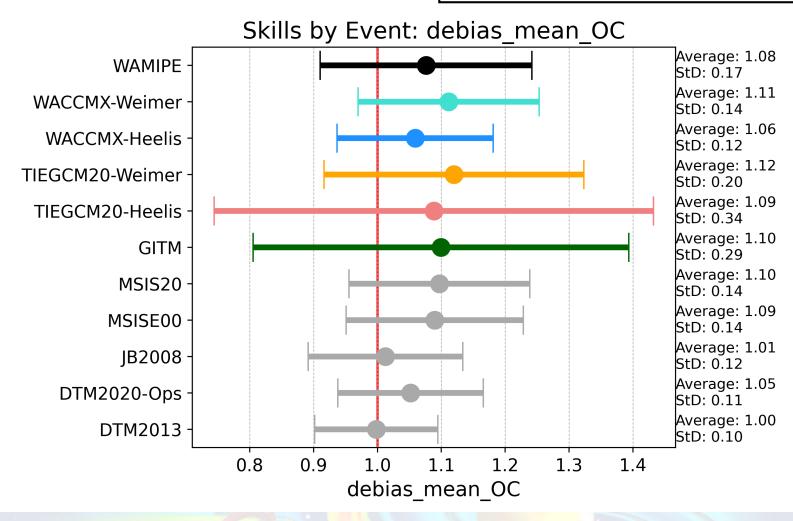


Validation of neutral density during 151 storm events from 2001 to 2023

O/C = 1, no model bias on average (best)

> 1, underestimate

< 1, overestimate

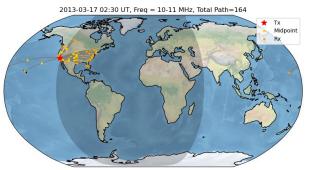


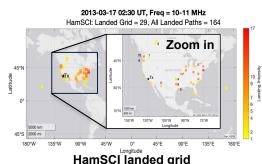


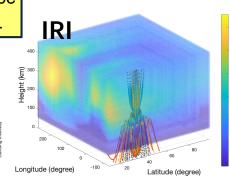


Ionosphere Model Validation with Amateur Radio Community (HamSCI), led by Dr. Min-Yang Chou

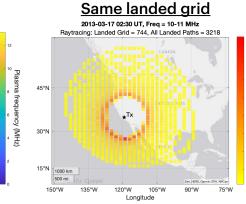
Goal: To reproduce the **HF propagation** from the **HamSCI data**, we will be able to use HamSCI data for ionospheric research and **model validation**.







Raytracing landed grid



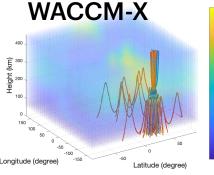
2013-03-17 02:30 UT, Freq = 10-11 MHz, [waccmx]
Raytracing: Landed Grid = 852, All Landed Paths = 3425

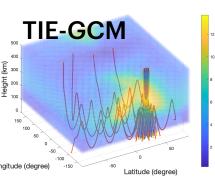
HamSCI

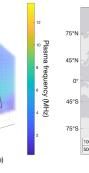
- An official NASA citizen science project
- allows researchers to collaborate with the amateur radio community in scientific investigations

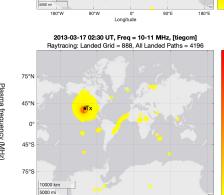
PHaRLAP (Ray tracing tool)

- Developed and controlled by Dr. Manuel Cervera, DoD, Australian Government.
- Matlab toolbox for modeling of the propagation of High Frequency (HF) radio waves
- Semi-open source (some of the key codes are precompiled.)





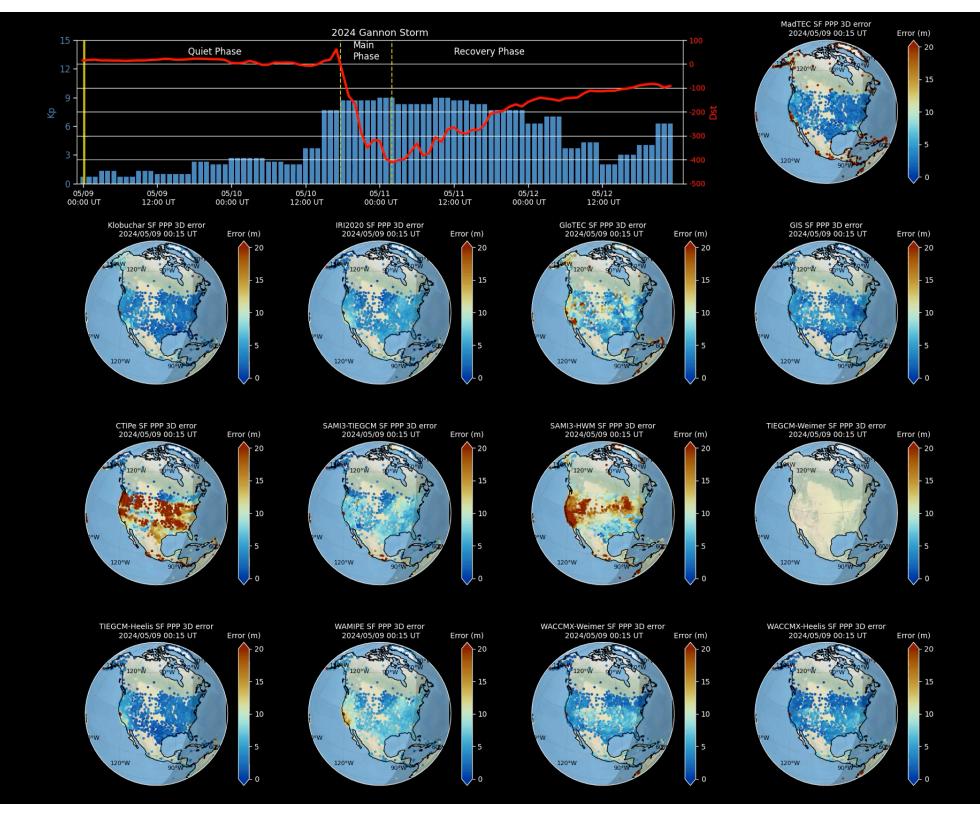




(not all raypaths)







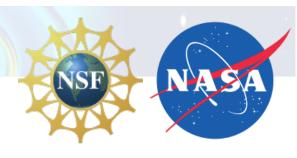
Future outlook

- Onboarding CESM3 as soon as it becomes available
 - 0.5-deg resolution, dynamical core to SE: resolving GWs is particularly important for plasma bubble seeding
- New NO scheme should improve neutral density
- SWMF/WACCM-X; WACCM-X/PBMOD
- Expand validation activities (e.g, satellite trajectories, see Jia Yue's science highlight talk on Thursday for an example)









NASA Community Coordinated Modeling Center (CCMC) WACCM-X/RoR Demo

Jack Wang and all CCMC members

Community Coordinated Modeling Center, NASA GSFC

June 25nd, 2025



Community Coordinated Modeling Center

About ~

Models 🗸

Simulation Services ^

Validation ~

Community Support 🗸

Execute Runs-On-Request

Search Runs-on-Request

Instant Run

Continuous/Realtime Run

Visualization

Publications Policy

Publications List

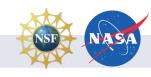
Change Log

Reau more



The TIE-GCM v3.0 mode community through the (ROR) service.







Home > CCMC Developed Web Apps/Tools

CCMC Runs on Request System

Runs on Request (ROR) is a simulation service accessible to anyone wishing to execute space science and space weather models hosted by the CCMC.

View all models available for Runs on Request (ROR) →

Track current active request progress on the ROR Requests Dashboard \rightarrow

Search ROR Archive →





Home

Model Catalog

WACCMX

Found 2 models

Reset Filters

WACCMX

Whole Atmosphere Community Climate Model With Thermosphere and Ionosphere Extension

Version: 2.2

Status: Production

→ R

Runs-on-Request

Public Repository





WACCM-X RoR submission interface

E-field model

- Weimer
- Heelis

Solar EUV model

- **EUVAC**
- FISM2 daily band

Request an WACCMX Model Run Please complete the form below. * Indicates a required field. **You must agree to the CCMC Data Policy in order to submit a run** Do you give your consent? **YES** Step 1: GENERATE YOUR REQUEST Your run results will be published online under your Run Registration Number (FirstName_LastName_MMDDYY_ModelType_RunNumber) WORK OR SCHOOL EMAIL Enter a valid work or school email how to contact you address FIRST NAME (GIVEN)* Enter your given name your given name LAST NAME (FAMILY)* Enter your family name your family name Unless you want to overwrite it **RUN NUMBER*** choose run number for today \$ max 15 runs per day **KEYWORD/S*** helps to sort and search the results of Enter a keyword simulations Step 2: Dates Range Select start and end dates of your simulation spanning at least 1 and at most 10 days. Start date Select or type a valid date in the YYYY/MM/DD suggested format, e.g. 2016/01/15 **End date** Select or type a valid date in the YYYY/MM/DD suggested format, e.g. 2016/01/15 Choose E-Field model Weimer Parameters of the solar and E-field Please, select a valid model for inputs are based on the observational the selected dates range or Weimer: 1996/01/01 - 2024/06/01 adjust the dates

Step 3: E-Field Model

dataset

Heelis: 1960/01/01 - 2024/06/01

Step 4: Solar EUV model

Choose EUV model

EUVAC (default)

A Tutorial on the NSF NCAR WACCI



Community Coordinated Modeling Center

About ~

Models 🗸

Simulation Services ^

Validation ~

Community Support 🗸

Execute Runs-On-Request

Search Runs-on-Request

News: TIE-GCM v

The TIE-GCM v3.0 mode community through the (ROR) service.

Instant Run

Continuous/Realtime Run

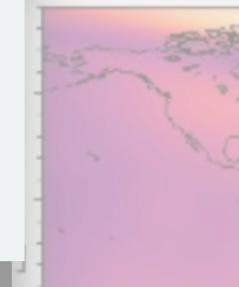
Visualization

Publications Policy

Publications List

Change Log

Reau more







Status	Run Number	Key Words	Model	Model Version	Validation Level	Year	DoY at Start	F10.7 at Start	F10.7 (three year average)	Run Type	Input Type	Event Date
running	Dr_SC_062325_IT_1	test_run	WACCMX	2.2		2024	131	0.00000	0.00	event		May 10, 2024
Published	Huang_Guo-wei_062025_IT_1	Guo	WACCMX	2.2		2024	151	0.00000	0.00	event	var	2024-05- 30
Published	Dmytro_Kotov_061425_IT_1	Ionospheric storms	WACCMX	2.2		2023	150	0.00000	0.00	event	var	2023-05- 30
Published	Dmytro_Kotov_061325_IT_4	Ionospheric storms	WACCMX	2.2		2023	150	0.00000	0.00	event	var	2023-05- 30
Published	Huang_Guo-wei_061325_IT_1	Guo	WACCMX	2.2		2024	142	0.00000	0.00	event	var	2024-05-
Published	Shrivansh_Panwar_060825_IT_1	Meridional Wind	WACCMX	2.2		2004	203	0.00000	0.00	event	var	2004-07-
Published	Huang_Guo-wei_060525_IT_1	Guo	WACCMX	2.2		2024	132	0.00000	0.00	event	var	2024-05- 11
running	Huang_Guo-wei_060425_IT_1	Guo	WACCMX	2.2		2024	132	0.00000	0.00	event		May 11, 2024
Published	Swati_Chowdhury_060225_IT_1	meridional wind	WACCMX	2.2		2020	337	0.00000	0.00	event	var	2020-12- 02
Published	Huang_Guo-wei_052825_IT_1	Guo	WACCMX	2.2		2024	122	0.00000	0.00	event	var	2024-05- 01
Published	chaneMoges_Seid_052725_IT_1	temperature	WACCMX	2.2		2019	339	0.00000	0.00	event	var	2019-12- 05
Published	Meenakshi_S_052625_IT_2	wave4	WACCMX	2.2		2020	14	0.00000	0.00	event	var	2020-01- 14
running	Huang_Guo-wei_052325_IT_1	Guo	WACCMX	2.2		2024	122	0.00000	0.00	event		May 01, 2024
Published	Sophia_Laranja_052325_IT_1	WACCMX_10032014_Weimer_EUVAC	WACCMX	2.2		2014	276	0.00000	0.00	event	var	2014-10-
Published	Sophia_Laranja_052325_IT_3	WACCMX_11202013_Weimer_EUVAC	WACCMX	2.2		2013	324	0.00000	0.00	event	var	2013-11-
Published	Sophia_Laranja_052325_IT_4	WACCMX_11202013_Heelis_EUVAC	WACCMX	2.2		2013	324	0.00000	0.00	event	var	2013-11-
Published	Sophia_Laranja_052325_IT_2	WACCMX_10032014_Heelis_EUVAC	WACCMX	2.2		2014	276	0.00000	0.00	event	var	2014-10-
Published	Sophia_Laranja_052225_IT_3	WACCMX_01172012_Heelis_EUVAC	WACCMX	2.2		2012	17	0.00000	0.00	event	var	2012-01- 17
Published	Min-yang_chou_052225_IT_1	gannon storm for sami3	WACCMX	2.2		2024	130	0.00000	0.00	event	var	2024-05- 09

Fatemeh_Bagheri_042225_IT_1

Run Status: Run Complete

Status updated: 2025-04-24T21:40:58+0000

Run Metadata

Metadata Record:	View Full Run Metadata in the CCMC Metadata Registry (CMR)
Metadata as JSON:	View Full Run Metadata as JSON
Model Domain:	IT
Model Name:	WACCMX
Model Version:	2.2
Key Word:	tides

CS output:	GEO
Run type:	event
Boundary condition type:	var
Year run:	2017
DOY:	272
Start time:	2017/09/29 00:00:00
End time:	2017/09/30 00:00:00
E-field model:	weimer

Output Data

- View 3D Ionosphere/Thermosphere
- Create Timeseries in 3D Ionosphere/Thermosphere

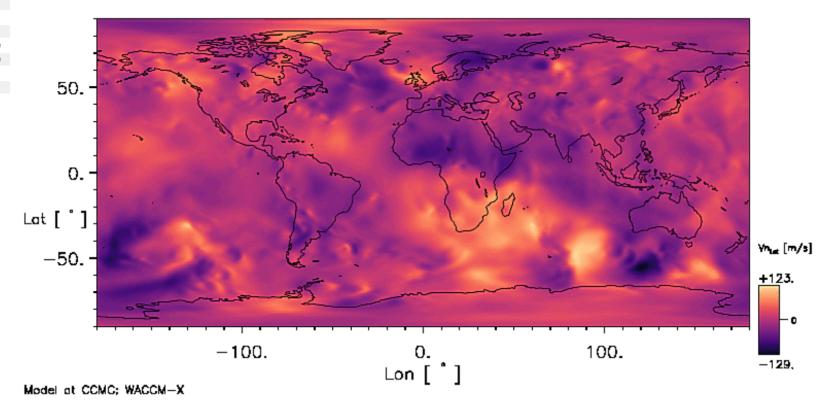
Run Services

- Request output data in bulk
- Browse individual output files

Supplementary Services

• CCMC DONKI notifications during the run period:

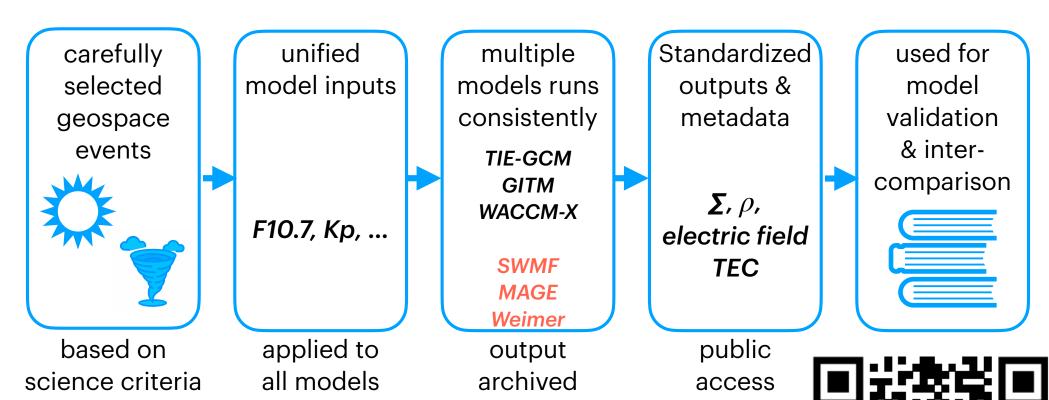
10/01/2017 Time = 00:00:00 UT IP= 64.50







Time Period database: a curated collection for consistent model evaluation



All model simulation output are publicly available on CCMC TP database.

(https://kauai.ccmc.gsfc.nasa.gov/CMR/ TimeInterval/viewAllTI)





Back-ups





Space weather societal impact

- Ionosphere variability (navigation, communications)
- Atmosphere variability (satellite/debris drag)
- Geomagnetically induced currents GICs (electric power systems)
- Near-earth radiation and plasma environment (aerospace assets functionality)
- Solar energetic particles SEPs (human exploration, aviation safety, aerospace assets functionality)
- Galactic cosmic rays GCRs (human exploration, aviation safety, aerospace assets functionality)





TEC 2013 March Storm, Lat x LT. map

